# A numerical study of the Trojan dynamics

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Collaborations with:

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### Restricted "several" bodies-problem

Different models:

asteroid + Sun + 8 planets asteroid + Sun + 4 giant planets asteroid + Sun + Jup + Sat

From 
$$3+8*3 = 24$$
 to  $3+2*3 = 9$   
d. f.

But even 9 D.F. imply numerical studies: num. integrations of the trajectories +

Analysis of the Traj.: Lypunov exponents, Fourier analysis, Frequency Analysis...





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 $(\lambda, \varpi, \Omega)$  Fixed



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#### Overlap of MMR above coll. lines: Global chaos

#### Projection of the observed inner solar system's objects on the ecliptic



Main asteroids belt (~400000)

Terrestrial planet's crossers (~5000) Comets

(~200)

Jupiter's trojans (~2000)

Trojans can orbit far from L4 or L5 (2 to 2.5 A.U.)

### Kepler (1609)



- $oldsymbol{\omega}$  : argument of the perihelion
- e : eccentricity
- M : mean anomaly

 $\lambda=M+arpi$  : mean longitude

a : semi-major axis

 $\varpi=\omega+\Omega~$  : longitude of the perihelion

~ actions 
$$\left( egin{array}{ccc} a, \ e, \ i, \end{array} & \mathbf{M} = nt \end{array} 
ight.$$
 ~ angles  $\mathbf{M}, \ \mathbf{\omega}, \ \mathbf{\Omega} \end{array} 
ight)$ 



 $\varOmega$  : longitude of the node

i : inclination

## fundamental Frequencies (proper frequencies)

3 frequencies for the Trojan: (n, g, s)

Orbital motions: periods 12 years for Jupiter 164 years for Neptune 1000 years at 100 A.U.

Secular motions: periods > 25000 years

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## fundamental Frequencies (proper frequencies)

#### If Q.P. solution (evolves on a KAM torus)

3n-I planetary frequencies:  $(n_j, g_j, s_j)$  one of the  $s_j = 0$ 3 frequencies for the Trojan: (n, g, s)

	3 months for Mercury	
periods	12 years for Jupiter	
periods	164 years for Neptune	
	1000 years at 100 A.U.	
	periods	

Secular motions: periods > 25000 years

 $H(I,\theta) = H_0(I) + \varepsilon H_1(I,\theta)$  H real analytic for  $(I,\theta) \in B^n \times \mathbb{T}^n$ 

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If  $\varepsilon = 0$ 

$$F: B^n \longmapsto \Omega \subset \mathbb{R}^n$$
$$I \longmapsto \nu(I) = \nabla H_0(I)$$

if 
$$det\left(\frac{\partial^2 H_0(I)}{\partial I^2}\right) \neq 0$$

F is a diffeo. (loc.)

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There exists  $\Omega_{\varepsilon}$  set of diophanine frequiencies  $\leftrightarrow$  KAM tori

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Pöschel (1982),: There exists a diffeo.  $\Psi$  and a coord. syst.  $(\varphi, \nu)$  such that  $\Psi: \mathbb{T}^n \times \Omega \longrightarrow \mathbb{T}^n \times B^n$   $\Psi$  is analytical/ $\varphi$  and  $C^{\infty} / \nu$  $(\varphi, \nu) \longmapsto (\theta, I)$  The flow is linear on:  $\mathbb{T}^n \times \Omega_{\varepsilon}$ :  $\dot{\nu} = 0$ ,  $\dot{\varphi} = \nu$ 

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For fix 
$$\theta \in \mathbb{T}^n$$
:  $\theta = \theta_0$   
 $F_{\theta_0} : B^n \longrightarrow \Omega$ ;  $I \longrightarrow p_2(\Psi^{-1}(\theta_0, I))$ 

The frequency map  $F_{\theta_0}$  is a smooth diffeo. from, the actions space to the frequencies space

### Goal

to obtain numerically a frequency map:

defined on  $B^n$ 

which coincide with  $F_{\theta_0}$ , up to numerical accuracy, on the set of KAM tori

numerical tool: Frequency analysis (J.Laskar, 1988, 1990)

Quasi-periodic decomposition of  $ae^{i\lambda} = \sum_k \alpha_k e^{if_k}$ 

$\begin{vmatrix}  \alpha_j  \\ (AU) \end{vmatrix}$	$f_j$ rad/yr	combinations
46.183882	.02005033	n
.259757	.52968580	$n_5$
.058931	.21330868	$n_6$
.049411	.02004870	$n-g_8+g$
.040885	.02005196	$n+g_8-g$
.038045	.01808704	$-n+n_8+g$
.031431	.02201360	$3n - n_8 - g$
• • •	• •	

Quasi-periodic decomposition of  $z_5 = e_5 \exp i \varpi_5$ 

$$z_5(t) \approx \sum_{j=1}^N \alpha_j \exp\left(if_j t\right)$$

$$f_j = k_5 n_5 + k_6 n_6 
 p_5 g_5 + p_6 g_6 + q_6 s_6$$

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$ lpha_j $	$f_j$ ("/yr)	$k_5$	$k_6$	$p_5$	$p_6$
$4.41 \times 10^{-2}$	$+4.027603 \times 10^{0}$	+0	+0	+1	+0
$1.59 \times 10^{-2}$	$+2.800657  imes 10^{1}$	+0	+0	+0	+1
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$$f_j = k_5 n_5 + k_6 n_6$$
$$p_5 g_5 + p_6 g_6 + q_6 s_6$$

# **Dynamical Maps and Frequency Analysis**

C.I. planets

$$(a_j, e_j, I_j, \lambda_j, \varpi_j, \Omega_j)$$
 Fixed

$$(n_j,g_j,s_j)$$
 given













# Frequency Map (secular)



## Main secular resonances in the asteroid belt



$$pg + qs + r_1g_5 + r_2g_6 + r_3s_6 = 0$$

Label	$\mid p$	q	$r_1$	$r_2$	$r_3$
1	1	0	0	-1	0
2	1	0	-1	0	0
3	0	1	0	0	-1
4	1	1	-1	0	-1
5	1	1	0	-1	-1
6	1	0	1	-2	0
7	0	1	-1	1	-1
8	2	-2	0	0	0
9	1	-1	-1	0	1
10	1	-1	0	-1	1
11	2	1	0	-2	1
12	1	-2	0	-1	2
13	1	0	2	-3	0
14	1	-1	1	-2	1
15	1	-3	0	-1	3
16	1	-3	-1	0	3
17	1	-4	0	-1	4
18	2	-3	0	-2	3
19	1	-4	-1	0	4





### Jovian Trojans



Jupiter's trojans (~2000) Robutel, Gabern & Jorba (2005, 2006)



Robutel, Gabern & Jorba (2005, 2006)



## Comparison of different models



#### R.4.B.P. (Sun+J+Sat+T)

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E.R.T.B.P. (Sun+J+T)

### Comparison of different models


- 5 planetary frequencies :  $(n_5, n_6, g_5, g_6, s_6)$ 
  - 3 for a test-particle : (n,g,s)

Trojan : I:I orbital resonance  $n_5 = n$ 

(
u, g, s)

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 $u \in [7500,9200]$  arcsec/year  $\ T_{
u} \in [140,155]$  years

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 arcsec/year  $T_{v} \in [140,155]$  years $g \in [230,450]$  arcsec/year  $T_{g} \in [2880,5634]$  years $s \in [-50,5]$  arcsec/year  $T_{s} > 25000$  years

# 2 obvious families of resonances

Family I  $p\nu = n_5$ 



E.R.T.B.P. (Sun+J+T) 
$$p\nu - n_5 + qg = 0$$

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Family III: Secular resonances

 $kg + ls + k_5g_5 + k_6g_6 + l_6s_6 = 0$ 

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$$n_5 - 2n_6$$

$$\frac{2}{5}(n_5 - 2n_6) \approx 8500"/yr$$
$$\nu \in [7500, 9200]"/yr$$

Family II

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$$2n_{5} - 5n_{6} (+2g_{6})$$

$$\frac{2}{5}(n_{5} - 2n_{6}) \approx 8500^{\circ}/yr$$

$$\nu \in [7500, 9200]^{\circ}/yr$$
Family II

$$\frac{2n_5 - 5n_6}{4} \approx 350"/yr$$

$$g \in [230, 450]"/yr$$
Family IV

Secondary resonances

Famille I:  $pv - n_5 + qg + q_5g_5 + q_6g_6 = 0$ 

Famille II:  $5v - 2(n_5 - 2n_6) + pg + p_5g_5 + p_6g_6 = 0$ 

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Secular resonances

Famille III:  $q s + q_6 s_6 + p_5 g_5 + p_6 g_6 = 0$ 

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#### G. I. + secular frequencies



### E.R.T.B.P (S+J+T)





$$12v - n_5 + qg = 0$$
 avec  $q \in \{8, \dots, 13\}$   
 $13v - n_5 + qg = 0$  avec  $q \in \{-4, \dots, 8\}$   
 $14v - n_5 + qg = 0$  avec  $q \in \{-3, \dots, 3\}$ 



















$$13\nu - n_{5} + qg + q_{5}g_{5} + q_{6}g_{6} = 0$$

$$14\nu - n_{5} + qg + q_{5}g_{5} + q_{6}g_{6} = 0$$

$$14\nu - n_{5} + qg + q_{5}g_{5} + q_{6}g_{6} = 0$$

$$5.25 \quad 5.30 \quad 5.35 \quad 5.40$$

$$5.25 \quad 5.50 \quad 5.35 \quad 5.40$$

$$5\nu - 2(n_{5} - 2n_{6}) - 0g + p_{5}g_{5} + p_{6}g_{6} = 0$$

$$5\nu - 2(n_{5} - 2n_{6}) - 1g + p_{5}g_{5} + p_{6}g_{6} = 0$$

$$5\nu - 2(n_{5} - 2n_{6}) - 2g + p_{5}g_{5} + p_{6}g_{6} = 0$$

# Long-term stability






















 $s = s_6$ 



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Regions where  $\sigma > -3 \;$  (orange, red) are cleared in IGy except 2



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Gap along  $4g + (2n_5 - 5n_6) - g_6 = 0$ 



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 $s = s_6$ 

Regions where  $\sigma > -3$  (orange, red) are cleared in IGy except 2

Gap along  $4g + (2n_5 - 5n_6) - g_6 = 0$ 



Slow diffusion along

$$4g + (2n_5 - 5n_6) - g_5 = 0$$





Slow diffusion along

$$4g + (2n_5 - 5n_6) - g_5 = 0$$

during 600 My

then wandering for 200 Ma

and ejection at 800 My



## Overlapping in family II

$$5v - 2(n_5 - 2n_6) - 0g + qg_5 - (q+2)g_6 = 0$$

## bounded diffusion







